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THE
THEORY
OF
HADLEY's QUADRANT,
OR
THE RULES FOR THE CONSTRUCTION
AND
USE OF THAT INSTRUMENT
DEMONSTRATED.

By the Reverend Mr. LUDLAM.

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ADVERTISEMENT.

*T*H E following sheets are designed as a Supplement to the Directions for the Use of HADLEY'S QUADRANT, &c. That tract is intended to teach the practical part, this to show the theory. Each part being thus printed separately, the maker and practical observer will not be perplexed with abstruse demonstrations in no wise necessary for them; nor the man of science find the reasoning interrupted and obscured by matters foreign to the theory,

AN ALMANAC FOR THE YEAR

BY JAMES COOPER
ENTITLED, THE NEW ENGLISH ALMANACK,
FOR THE YEAR OF CHRIST 1771.
CONTAINING A DETAILED ACCOUNT
OF THE SUN'S POSITION, AND
MOVEMENTS, FOR THE WHOLE YEAR,
WITH THE ECLIPSES, AND
MOVEMENTS OF THE MOON,
AND OTHER PLANETS,
THE EQUINOXES, ETC.

THE THEATRE OF HADLEY's

QUADRANGLE &c. STAMPED

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1. LET NIn , and RHr , fig. 1. represent the sections of the plane of the figure by the surfaces of two speculums standing thereon. Let BI be a ray of light falling on a point I of the first speculum, and reflected thence to a point H of the second speculum, and by it into the line HO ; then the angle between the incident ray, and the ray after two reflections, is double the angle between the speculums.

2. Since the incident ray BI is in the plane of the figure, the reflected ray will be so too *; and because the ray IH is in that plane, so will also the reflected ray HO *; the rays BI and HO therefore being both in the plane of the figure, let them (when produced) meet in E , and let the two sections NIn and RHr (produced) meet in S , then will ISH be the inclination of the reflecting planes to each other †; but ISH is the difference of the angles IHR and

* Smith's Optics, Art. 7. † Def. 6. and Prop. 19. El. XI.

THEORY OF

*HIS**. Produce *EH* backwards to *A*, and because by the laws of reflection \dagger $IHR = EHS = RHA \ddagger$, therefore *IHA* is double of *IHR*. In like manner because $\dagger HIS = BIN = SIE \ddagger$, therefore *HIE* is double of *HIS*; but *IEH* is the difference between *IHA* and *HIE**, that is between twice *IHR* and twice *HIS*, or it is double the difference of the angles *IHR* and *HIS*, that is double the angle *ISH*.

3. This demonstration is general, but the figure represents the disposition of the glasses and course of the rays in the fore observation. It may be more convincing, and help us to clearer notions, to see it particularly applied to the back observation. Fig. 2 then represents the disposition of the glasses and course of the rays in the back observation. In this let the two sections *NIn* and *RHr*, as before, meet in *S*. Produce *HS* to *T*; and *IST* is the inclination of the reflecting planes to each other \parallel . Let *BI* the incident ray, and *HO* the ray after two reflections (when produced) meet in *E*; and produce *OHE* to *A*, so that *IEA* may be

* 32 El. I. † Smith's Optics, Art. 9. ‡ 15 El. I.
|| Def. 6. Prop. 19. El. XI.

the



the obtuse angle which these rays make with each other. Then is IEA the sum of IHE and HIE *, the former of which (by the laws of reflection †) is double of IHS , and the latter double of HIS . Therefore IEA is double the sum of IHS and HIS ; that is double of IST *.

4. Let SH and SI , fig. 3. represent the sections of the plane of the figure by the surfaces of two speculums standing perpendicularly thereon. Let these sections meet in S , and ISH will be the inclination of the reflecting planes, and S the point where their common section passes the plane of the figure, to which it is also perpendicular ‡. Let Q be an object placed between the speculums, and in the plane of the figure. Several representations of that object will be formed by the successive reflections of these speculums; all which will lie in a plane passing through the object, and perpendicular to both the reflecting planes; that is in the plane of the figure as was before shewn, par. 2. || But these representations

* 32 El. I. † Smith's Optics, Art. 9, &c. as before.
 ‡ Def. 6. and Prop. 19. El. XI. || Smith's Optics,
 Art. 7, &c.

or images of the object Q , will also lie in the circumference of a circle drawn on that plane, whose center is S , and radius SQ the perpendicular distance of the object from the common section of the two speculums.

5. Draw QT perpendicular to the speculum SH , produce QT to A , so that $TA=TQ$ and Q being considered as the focus of rays incident on SH , A will be the focus of reflected rays *, or place of the image after one reflection from SH †. But the right-angled triangles STS and ATS having the sides TA and TQ equal, and ST common, will have their hypothenuses SQ and SA equal ‡. Therefore the circle described with the center S and radius SQ , will pass through A . — Moreover the arch HA is equal to HQ .

6. For the like reason if a be the image made by one reflection from the other speculum SI , it will lie in the circumference of this circle passing through Q ; and the arch Ia is equal to IQ .

7. The image a may be considered as an object in respect of the speculum SH ||.

* Smith's Optics, Art. 202.

† Ibid, Art. 25.

‡ 4 El. 1.

|| Smith's Optics, Art. 111.

Take

Take therefore the arch $HB=Ha$, and B will be the image of that image, or the image seen in the speculum SH by two reflections; first from SI , and then from SH *.

8. The image A may in like manner be considered as an object in respect of the speculum SI †. Take the arch $Ib=IA$, and b is the other image made by two reflections and seen in the speculum SI .

9. Again take $HC=Hb$, and C will be the image seen in the speculum SH after three reflections; first from SH , then from SI , and then from SH again.

10. Take $Ic=IB$, and c is the other image seen (in SI) after three reflections.

11. Again take $HD=Hc$, and D is the image seen in SH after four reflections.

Take $Id=IC$, and d is the other image seen after four reflections.

12. Again take $HE=Hd$, and E is the image seen in SH after five reflections; and so on.

13. Cor. I. Since $Qg=2Ia$, and $Ba=2Ha$, subtract the former from the latter, and there

* Par. 5.

† Smith's Optics, Art. 111.

remains

remains $QB = 2 \times \overline{Ha - Ia} = 2HI$. Again, $Bc = 2Ic$ and $Dc = 2Hc$; subtract the former from the latter, and there remains $BD = 2 \times \overline{Hc - Ic} = 2HI$, and so on. Make therefore the arches QB , BD , DF , &c. each equal to $2HI$; and B , D , F , &c. will be the successive images of the object Q after an even number of reflections.

14. In like manner take Qb , bd , df , each equal to $2HI$; and b , d , f , &c. will be the place of the images seen in the other speculum after an even number of reflections.

15. Cor. 2. The distances of each of the images B , D , F , &c. or of b , d , f , &c. from the object Q , measured along the arch of the circle; is equal to the arch HI multiplied by the number of reflections.

16. Cor. 3. Again; since $HA = HQ$, and $HC = Hb$, take the former from the latter, and there remains $AC = Qb = 2HI$ *. Therefore if QA be taken equal to $2QH$, and the several arches AC , CE , &c. each equal to $2HI$; then A , C , E , &c. will be the place of the images seen in SH after an odd number of reflections.

* Par. 14.

17. In

17. In like manner take $2g = 2QJ$, and then the several arches ac , ce , &c. each equal to $2HI$ and a , c , e , will be the place of the images seen in the other speculum after an odd number of reflections.

18. *Cor. 4.* The distance of the first image from the object or QA , is equal to $2QH$: that of the third image or QC ($= QA + AC$) is equal to $2QH + 2HI$; that of the fifth image (or $QA + AE$) is $2QH + 4HI$, &c. the distance of each image from the object Q , measured along the arch of the circle, being equal to the arch HI multiplied by the number of reflections save one, and increased by twice QH the distance of the object from the speculum.

19. The same is true of a , c , e , &c. the images seen in the other speculum after an odd number of reflections.

20. *Cor. 5.* If the two speculums, keeping their inclination to each other, be turned round their common section as an axis; the images B , D , F , &c. formed by an even number of reflections will be immovable; because the intercepted arch HI is not altered: but the alternate images A , C , E , &c. formed by an odd number of reflections,

will

will change their places by double the change of the arch QH , or double the angular motion of the speculums.

21. *Cor. 6.* Hence (in fig. 21) with the center S and radius SB describe a circle cutting in A , the last reflected ray HO produced backwards; and A will be the place of the image of the object B after both reflections. For all the images of this object lie in the circumference of that circle *, and the last image must lie in the direction of the last reflected ray †; therefore it is in A their common intersection.

22. Let an object be placed at C in the reflected ray OH produced, so as to be seen by direct vision through the speculum H along with the reflected image A . The image A will continue immoveable notwithstanding the two speculums be turned together round their common intersection as an axis ‡; and therefore will seem always to adhere to the object C .

23. Let the eye be placed at E the intersection of the incident ray with the ray after two reflections; then BEA , or BEC

* Par. 4 and following.
102.

† Smith's Optics, Art. 101,
Par. 20.

the

the angle which these two rays make with each other, is also the angle which the object *B* makes with its image *A*: it is also the angle which the two objects *B* and *C* subtend to the naked eye at *E*. This angle *BEA* has been shown to be equal to twice *ISH* *.

24. If the eye be placed any where else in *HE* produced as at *O*, the doubly reflected image *A* and the object *C* will still appear to coincide as before: but the angle which the objects *B* and *C* subtend to the naked eye is now *BOC*, differing from *BEA* by the angle *EBO*, or *IBO* †.

25. Let now the speculum *NIn* be turned about the point *I*, so that the doubly reflected image of the object *G* (placed in the line *HO* produced) may be made to coincide with that object itself seen directly by the eye at *O*; then is the angle *ICH* double of *ISH* the angle between the speculums. For in this case the incident ray is *CI*, and the ray after two reflections is *HO*, which produced backwards passes through *C* by the supposition of the coincidence aforesaid. Therefore the inter-

* Par. 2.

† 32 El. I.

section of these two rays is at *C* the object itself, and consequently *ICH* the angle of their intersection is double of *ISH* the angle of the speculums *.

26. The intersection before marked *E*, in this case removes to *C* on the other side of *H*; and *S*, the point of intersection of the two sections *NIn* and *RHr*, falls on the other side of the line *IH* †.

27. If

* Par. 2.

† To consider the motion of the points *E* and *S* more particularly, imagine the speculum *NIn* to turn round the point *I*, while the other speculum *RHr*, and likewise the position of the ray *HO*, remains fixed.

And first let the speculum *NIn* turn so that the point *S* may continually recede from *H*. Then the point *E* will continually recede from *H* towards *O*. When the two speculums become parallel, the point *S* and likewise the point *E* will go off *in infinitum*, as is evident because *BEA* is twice *ISH*; therefore when the latter vanishes, so does the former. In this case the incident ray and ray after two reflections are parallel.

Continue the motion of the speculum *NIn*, and the sections *NIn* and *RHr* will now meet on the other side, in *SH* continued. The intersection of the incident ray and ray after two reflections will now also remove to the other side, in *OH* continued. This will appear, if instead of considering *HO* as the ray after two reflections, we consider it as the incident ray, and therefore *HI* as a ray given in position and incident on the second speculum *NIn*; *IC* will then be the ray after two reflections by Smith's Optics, Art. 11.

Continue the motion of the speculum *NIn* and the point of intersection of these two rays will continue its motion from *A* towards *H*, and passing the point *H* will recede beyond *O* *in infinitum*. Thus these two rays will again be parallel. This happens when the speculum *NIn* makes a right angle with the speculum *RHr*.

Continue this motion, and the intersection of the incident ray and ray after two reflections will a second time remove

27. If the object *C* be removed to such a distance that the angle *ICH* vanishes, the angle *ISH* will also vanish, and the speculums become parallel to each other.

The incident ray *CI*, and the ray after two reflections *HO*, in this case are also parallel, their intersection *C* being removed to an infinite distance.

28. *Scholium.* If when the index stands at the beginning of the divisions, the two speculums are put into a parallel position, by turning one of them about till the reflected image of some distant object *C* coincides with that object seen directly, the instrument is then said to be *adjusted*. The instrument being thus adjusted and the speculums fixed, the index will always show

to the other side, in *OH* continued. When the face of the speculum comes into the direction *IH*, the intersection of those two rays will then be in *H*. After this the back of the speculum will be turned towards *RHr* for an intire semi-revolution, till the face again comes into the direction *IH*, and the two rays aforesaid again intersect in *H*; the speculums then making an acute angle the same way as at first.

If the motion of the speculums be still continued, the intersection of the two rays at *E* will recede from *H* towards *O* *in infinitum*, till the speculums are parallel as before.

The incident ray and ray after two reflections are twice parallel in one semi-revolution of the speculum *NIn*; viz. when the two speculums are parallel, and also when they are perpendicular to each other. In the other semi-revolution, the back of the speculum *NIn* is turned towards the other speculum *RHr*.

the double inclination of the speculums, either on the quadrantal arch, or arch of excess, according as the point *S* falls on one side or the other of the perpendicular let fall from *I* on *RHr* produced.

If, when the speculums are put into a parallel position as before, the index points to any number of degrees or minutes on the quadrantal arch (instead of 0) then it will always give the double inclination too much on the quadrantal arch, and equally too little on the arch of excess, by the number of degrees and minutes shown as before; and contrariwise.

29. The angle *BIC*, which any two objects *B* and *C* subtend at *I* the center of the index glass, is equal to the sum of the two angles *BEC* and *ICE* *.

30. Scholium 2. The quadrant being adjusted as before by a distant object, the former of these angles *BEC* will be shown on the quadrantal arch when the reflected image of *B* coincides with *C* seen directly. The latter of these angles *ICE* will be shown on the arch of excess when the reflected image of *C* coincides with that object itself seen

* 32 El. I,

directly,

directly. Or the sum of these two angles will be shown at once on the quadrantal arch, if, when the object *C* and its image coincide, the index be made to point at the beginning of the divisions; and then be carried forward till the reflected image of *B* coincides with *C* seen directly.

31. In figure 4. let *I* be the index-glass, standing parallel to *H* the fore horizon glass; let *R* be the back horizon glass, standing at right angles to *H* and *I*. Draw *PI* and *PR* perpendiculars to *I* and *R* meeting each other in *P*. Let *QH* be perpendicular to *H*. Let *BI* be the ray incident on *I*, and reflected from thence into *IH*, from thence into *HO*, the axis of vision for the fore horizon glass. Let *IR* be the ray reflected from the index-glass to the back horizon glass, and from thence into *RY* the axis of vision for that glass. Then by the laws of reflection $PIH = PIB$ *. Moreover $PIH = IHQ$ †. Therefore the angle which the line, joining the centers of the index-glass and fore horizon glass, makes with a perpendicular to the index-glass when the index stands at the beginning of the divi-

* Smith's Optics, Art. 8.

† 29 El. I.

fions,

sions, is equal to the angle of incidence on the index-glass in that position; and this is the constant angle of incidence on the fore horizon glass.

32. That $PIR = PIH + HIR$ is self-evident, whatever be the position of the perpendicular PI . But PIH is the angle of reflection from the index-glass to the fore horizon glass, and PIR the angle of reflection to the back horizon glass, which are equal to their respective angles of incidence. Therefore the angle of incidence on the index-glass in the back observation, exceeds the angle of incidence in the fore observation (the perpendicular PI standing in the same place) by the angle HIR which the centers of the two lesser mirrors subtend at the center of the great one.

33. When I and H are parallel, PIH is equal to IHQ , and IPR is a right angle. Therefore PRI is the complement of PIR , or of $PIH + HIR$, that is of $IHQ + HIR$. Therefore the angle of incidence on the back horizon glass, is constantly the complement to 90 degrees, of the sum of the angle of incidence on the fore horizon glass, and the angle which the centers of the two lesser

lesser mirrors subtend at the center of the great one.

34. Let RN represent the surface of the back horizon glass, RY the ray reflected from it or the axis of vision, and produce YR to M . Let OH be the axis of vision for the fore observation. Produce OH till it meets RY in Y , and let OH meet IR in Z . Then the angle RYZ , which the two axes of vision make with each other, is the difference of ZRM and YZR *. Now the former of these $ZRM = 2 \times ZRN + \dagger = 2 \times PIR \ddagger = 2 \times PIH + 2 \times HIZ$. The latter of those angles or $YZR = ZHI + HIZ * = \ddagger BIH + HIZ \ddagger = 2 \times PIH + HIZ \dagger$. Subtract the latter from the former, and we have their difference, or RYZ equal to HIZ , or HIR . Therefore the angle which the two axes of vision (for the fore and back observation) make with each other, is equal to that which the two lesser mirrors subtend at the center of the great one.

35. To show the correction necessary to be made in taking the height of a building by reflection from water, let figure 5 represent the position of the octant when the

* 32 El. I. † Smith's Optics, Art. 8. ‡ 29 El. I.
fore

fore observation is used; figures 6 and 7 show the position of the octant in the back observation. In these figures let LR be the surface of the water. BL the building, AL its reflected image seen in the water. R the point of reflection from the water. I the point of reflection from the index-glass. H the point of reflection from the horizon-glass, through which the extremity of the image A is seen in the water. O the place of the eye. E the intersection of the ray (from B the top of the building) incident on the index-glass with that ray after two reflections. S the point where the sections of the speculums by the plane of the octant meet each other; and continue HS to T in figures 6 and 7. Then is AEB the angle shown on the limb, being double of ISH in figure 5, and double of IST in figures 6 and 7 *, and BRL is the angle of the elevation of the building above the level of the water, supposing the eye placed at R the point of reflection. Now $2 \times BRL = BRA + AEB + EBR$ in figures 5 and 6, and equal to $AEB - EBR$ in figure 7 †; that is $2 \times BRL = AEB + IBR$ in figures 5

* Par. 1, &c. † Smith's Optics. Art. 8. ¶ 32 El. I.
and

and 6, and equal to $AEB - IBR$ in figure 7. Estimate therefore the angle IBR , which the middle of the index-glass, and the point of reflection from the water, makes at the top of the building: add this to the angle shown on the limb in the fore observation; add or subtract it in the back observation, according as the ray incident on the water passes without or within the index-glass; and the sum or difference gives the angle BRA ; half of which is the true angle of elevation.

36. To find the errors occasioned by a given deviation of the reflecting planes from their perpendicular position.

It may be convenient to consider the planes in question as the planes of several great circles of a sphere; and the point in which the two reflecting planes and plane of the octant all concur, as the center of that sphere. Their mutual inclinations may then be found by the common rules of spherical trigonometry.

37. In figure 8 let C be the center of the sphere, where the three planes before mentioned all meet. Let the plane of the great circle $TNEH$ represent the plane of the

D

octant.

octant. Let the plane of the great circle NCP represent the plane of the index-glass; and the plane of the great circle HCP , the plane of the horizon-glass. Then is NC and HC the intersections of the two reflecting planes with the plane of the octant, and PC their common intersection with each other. Moreover the spherical angles PNT and PHN are the inclinations of the reflecting planes to the plane of the octant which we here suppose equal; the spherical angle NPH is their inclination to each other. The arch NH measures the angle NCH , which the two sections of the reflecting planes by the plane of the octant, make with each other; and which is half the angle shown by the index on the limb. With the pole P describe a great circle $LDEF$, cutting the circle PND at right angles in D ; the circle PFH at right angles in F , and the circle $TNEH$ obliquely in E ; CE being the intersection of the plane of the oblique circle $LDEF$ with the plane of the circle $TNEH$. Through P and E draw a great circle; this will meet the circle $LDEF$ at right angles in E ; and the plane of this circle will pass through both



both PC the common intersection of the reflecting planes, and also through CE the intersection of the plane of the oblique circle $LDEF$ with the plane of the circle $TNEH$; the plane of this circle is seen edgewise in the figure. The spherical angle PEN which this circle makes with the circle TNE will be the inclination of PC the common intersection of the reflecting planes to the plane of the circle $TNEH$, that is to the plane of the octant, and NED will be the complement of this angle. Lastly, let the radius TC , lying in the plane of the great circle $TNEH$, represent a line drawn on the plane of the octant parallel to the axis of the telescope; passing both through CH the common section of the plane of the octant and horizon-glass, and through PC the common section of the reflecting planes. Through P and T draw the great circle PTL , meeting the circle TNE obliquely in T , but the circle $LDEF$ at right angles in L ; then will TCH be the (obtuse) angle which the line TC makes with the section CH . The arch PT is the measure of PCT , the angle which the line TC makes with PC the common section of the reflecting planes;

the arch TL is the complement of the arch PT , or it is the distance of the parallel circle from the great circle, mentioned by Hadley in his fifth corollary; that great circle being here represented by $LDEF^*$.

38. Fig. 9 represents the stereographic projection of these several circles on the plane of the circle $TNEH$ in figure 8, that is on the plane of the octant; the dotted lines CT , CN , CH , represent the several radii CT , CN , CH , lying in the plane of the circle $TNEH$. The several circles of the sphere before mentioned are all represented by circles in figure 9, according to the laws of the stereographic projection.

39. These things laid down, because the spherical angles PNT and PHN are by supposition equal, the arches EN and EH , likewise ED and EF are respectively equal: whence EN is equal to $\frac{1}{2}NH$; therefore EN is the measure of $\frac{1}{2}$ of the angle shown by the index on the limb. The arch $ED = \frac{1}{2}DF$, therefore ED is the measure of half the spherical angle NPH , or of half the true inclination of the reflecting planes to each other, or it is the measure of $\frac{1}{4}$

* See Philos. Transf. Nr 420.

of the corrected angle. — Now in the right-angled spherical triangle NDE we have the hypotenuse EN , and the angle $END = PNT$, to find the base ED . Therefore as radius is to sine END (or cosine of deviation from the perpendicular); so is sine EN to sine ED , which multiplied by 14 gives the corrected angle.

40. Next to find the inclination of PC the common section of the reflecting planes to the plane of the octant, or to find NED the complement of that angle. In the right-angled triangle NDE we have the hypotenuse EN and the angle END as before, to find the other angle NED . Therefore as radius is to tangent END , so is cosine EN to cotangent NED , or the tangent of the inclination of the common section of the reflecting planes to the plane of the octant.

41. Next to find PCT the angle which a line drawn on the plane of the octant, parallel to the axis of the telescope and passing through the common section of the reflecting planes, makes with that common section; or rather to find TL the arch which measures the complement of that angle.

First

First we have $ET = TH - HE = TH - EN$. Then in the right-angled spherical triangle TLE we have the hypotenuse ET , the angle $TEL = NED$ (just found) to find the perpendicular TL . Therefore as radius is to sine ET , so is sine TEL to sine TL , or distance of the parallel from the great circle in Hadley's fifth corollary.

42. Lastly by that corollary, as radius is to the sine complement of that distance, so is the chord of double the true inclination of the reflecting planes, or chord of the corrected angle; to the chord of the true angle between the observed objects: that is radius is to the sine complement of the distance aforesaid as the sine of half the corrected angle to the sine of half the true angle between the objects; the double of which is the angle required.

43. Of the effect of the refractions of the rays in a glass mirror whose two surfaces are parallel.

Let AL and BN in figure 10 be the two parallel surfaces of a glass mirror. Q the focus of incident rays. QC a ray incident on the first surface at C . Through Q draw QAB perpendicular to both surfaces, cut-

ting the first surface in *A*, the second in *B*, produce *AQ* backwards to *R*, so that *AR* may be to *AQ* as the sine of incidence to the sine of refraction, or as *m* to *n*, and *R* will be the focus after refraction at *C* *. Draw *CD* the refracted ray meeting the second surface in *D*. In the perpendicular *RS*, take *BS* equal to *BR*, and *S* will be the focus of rays reflected from the second surface at *D* †. Draw *DE* the reflected ray meeting the first surface in *E*. Take *AT* to *AS* as *n* to *m*, and *T* will be the focus of emergent rays after their last refraction at *E* ; that is, *T* will be the place of the last image formed by two refractions and one reflection. Take *AV* equal to *AQ*, and *V* will be the place of the image formed by the reflection of the first surface only, and *TV* will be the distance between these two images. I say that *TV* will be to *AB* the thickness of the glass, as twice the sine of refraction to the sine of incidence.

44. Produce *TE* to *M* so that *EM* may be the emergent ray; through *D* draw the perpendicular *DH*, and through *E* and *C* the perpendiculars *be* and *dc*. Produce *QC*

* Smith's Optics, Art. 223.

† Ibid. Art. 202, till

till it meets TE in F , and through F draw FG perpendicular to QT . Now $dCD = CDH^* = HDE + DEb^*$; but DC and DE may be considered as rays incident on AL ‡; and because the angles of incidence dCD and DEb are equal, therefore their angles of refraction cCQ and eEM are equal; but $cCQ = GQF^*$, and $eEM = GTF^*$, therefore the triangle QFT is isosceles and $QG = GT$. Now $AT = GT + AG$ and $AV = A\mathcal{Q} = G\mathcal{Q} - AG = GT - AG$; whence $AT - AV$ or $TV = GT + AG - GT - AG = 2 \times AG$. Now the triangles CFH and CQA , as also CDH and CRA are respectively similar: hence HF is to $A\mathcal{Q}$, also HD to AR , as CH to CA ; therefore HF is to HD as $A\mathcal{Q}$ to AR ||; that is, AG is to AB as $A\mathcal{Q}$ to AR , or as n to m ; consequently $2 \times AG$ or TV is to AB as $2n$ to m .

45. Cor. 1. If m be to n as 3 to 2; then $TV = \frac{4}{3}AB$ and $AG = \frac{2}{3}AB$ and $BG = \frac{1}{3}AB$.

46. Cor. 2. The last emergent ray proceeds exactly as if the reflection had been made without refraction, by a plane FG

* 29 El. I. † Smith's Optics, Art. 8. ‡ Ibid.
Art. 11. || 11 and 16 El. V.

parallel

parallel to BN the back surface, and distant from it $\frac{1}{3}$ of the thickness of the glass.

47. *Scholium.* The index-glass should be so fixed on the index; that the center on which the index turns may pass through the line FG ; that is, be distant from the back surface $\frac{1}{3}$ of the thickness of the glass; and be within it.

48. We shall here just mention the consequences that would follow from supposing the two surfaces of each glass mirror inclined to one another in a given angle: and we will consider first the effect of the refractions of the index-glass only.

49. When there is no refraction, or when the two surfaces of the index-glass are parallel, then the last image is removed from the object by twice the inclination of the reflecting surfaces of the two speculums, as we have shown before. But if the outer surface of the index-glass is inclined to the inner surface; the place of that image will then be changed by the refractions at the outer surface of the index-glass. The last image will in this case be formed by two refractions and one reflection at the index-

E glass:

glass: and its angular distance from the object will be increased, if the refracting angle of the index-glass be turned the same way with the angle which the speculums make with each other; and decreased if the refracting angle be turned the contrary way. This angular change in the place of the last image on account of refractions we will call *the angle of deviation.*

50. Now was the angle of deviation always the same in every position of the index-glass, the refractions then would occasion no error in the observed angle. For in this case the image, formed both by refraction and reflection, keeping one constant distance from that which would be formed by reflection only, will have the same angular motion, *viz.* double that of the index-glass. Therefore if any distant object and its last image, thus formed, be made to coincide when the index stands at the beginning of the divisions; the index will then always show the angular distance of this last image from the object, and consequently the angular distance of any other object, which is seen along with that last image; just as when there is no refraction. All the difference

will be, that the reflecting surface of the index-glass will not stand exactly parallel to the horizon-glass when the object and it's image are made to coincide, in order to adjust the instrument.

51. When the rays fall so nearly perpendicular on the index-glass, that the angles of incidence and refraction may be accounted to have the given ratio of their sines; then the angle of deviation is to the angle of inclination of the two surfaces of the index-glass, as double the difference between the sines of incidence and refraction is to the lesser of those sines. Hence the angle of deviation is nearly equal to the angle of the inclination of those surfaces.

52. But as those angles are not accurately as their sines, so the angle of deviation is not exactly what that rule gives it; nor the same in all positions of the index-glass. It is always greater than what the rule gives; but it is least of all when the ray within the glass falls perpendicularly on the reflecting surface, so that the first incident and last emergent rays coincide. This will be the case when *the index is in such a position, that the horizon-glass may be seen*

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by its own reflection. The incident ray in this case falls obliquely on the outer surface of the index-glass, and on that side of its perpendicular which is contrary to the refracting angle. We shall call this ray, *the ray of least deviation.*

53. The position of the index when the deviation is least being determined as before, the deviation will continually increase, as the distance of the index from this position increases on either side. If equal angular distances of the index from this position be taken on contrary sides, the angles of deviation in these two positions of the index will be nearly equal to each other, but not exactly so. That position of the index, in which the ray incident on the outer surface falls on the same side of the ray of least deviation with the perpendicular to that surface, will have the greatest angle of deviation.

54. All these consequences follow from the common properties of the prism, and need not be here particularly demonstrated. If a prism, whose refracting angle is about ten degrees, be properly placed before an index-glass whose surfaces are parallel, all these

these consequences may be shown experimentally *.

55. We have hitherto taken no notice of the refractions of the horizon-glass: it is because these refractions can never affect the observed angle. For as the horizon-glass is fixed, the rays always fall upon it with the same degree of obliquity. Hence the change in the place of the image, on account of the refractions of this glass, is constantly the same. The refractions of this glass may change a little the position of the axis of vision, but cannot alter the observed angle; not though the upper part of this glass be cut off, so that the object to be seen by direct vision may be unrefracted †, while

* It may be necessary to measure practically the angle of the prism. The following method is a very accurate one. Place the prism in a vertical position. Then, standing before the angle to be measured, place two candles or other proper objects, so that the image of one may be seen by reflection from one of the sides, and the image of the other be seen by reflection from the other of the two sides that form the refracting angle. Move these two objects till their reflected images coincide. Measure then the distances of these objects from each other and from the angle of the prism; from these distances compute the angle which the two objects subtend at the prism: half this is the angle sought. It may be necessary to cover the back of the prism, that the light refracted through the other two angles may not mix with that reflected from the sides.

† The angle of deviation from refraction only, is half that caused by refraction and reflection together as before.

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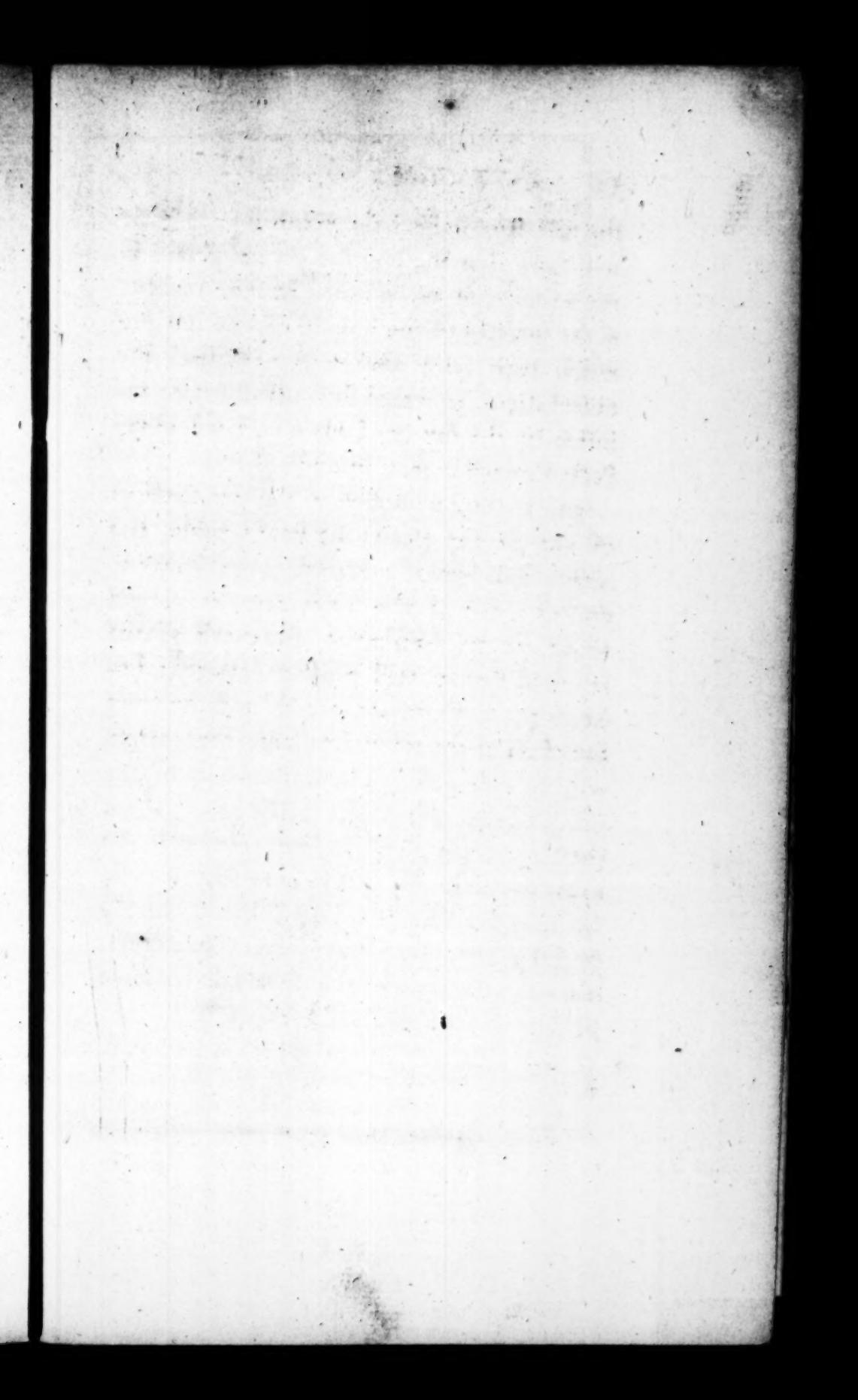
the refractions affect the other object. This will only alter the position of the glasses in the adjustment as before. In like manner if the surfaces of the dark glasses be not parallel, their refractions will not affect the observations, provided the instrument be adjusted by the sun, or some object whose reflected image is seen through them.

36. If the inclination of the two surfaces of any of the glasses be considerable, the object will appear coloured, or at least indistinct; and if the instrument be turned about (in its own plane) the place of the reflected image will now be changed; because the angle of incidence, and therefore the effect of the refractions upon that image, will be changed. The multiplicity of the images may also create confusion. Upon these accounts those surfaces ought to be made parallel. But if this was not the case, we may see that it is impossible to remedy perfectly the other errors from refraction, notwithstanding what has been said by some persons of distinguished learning *.

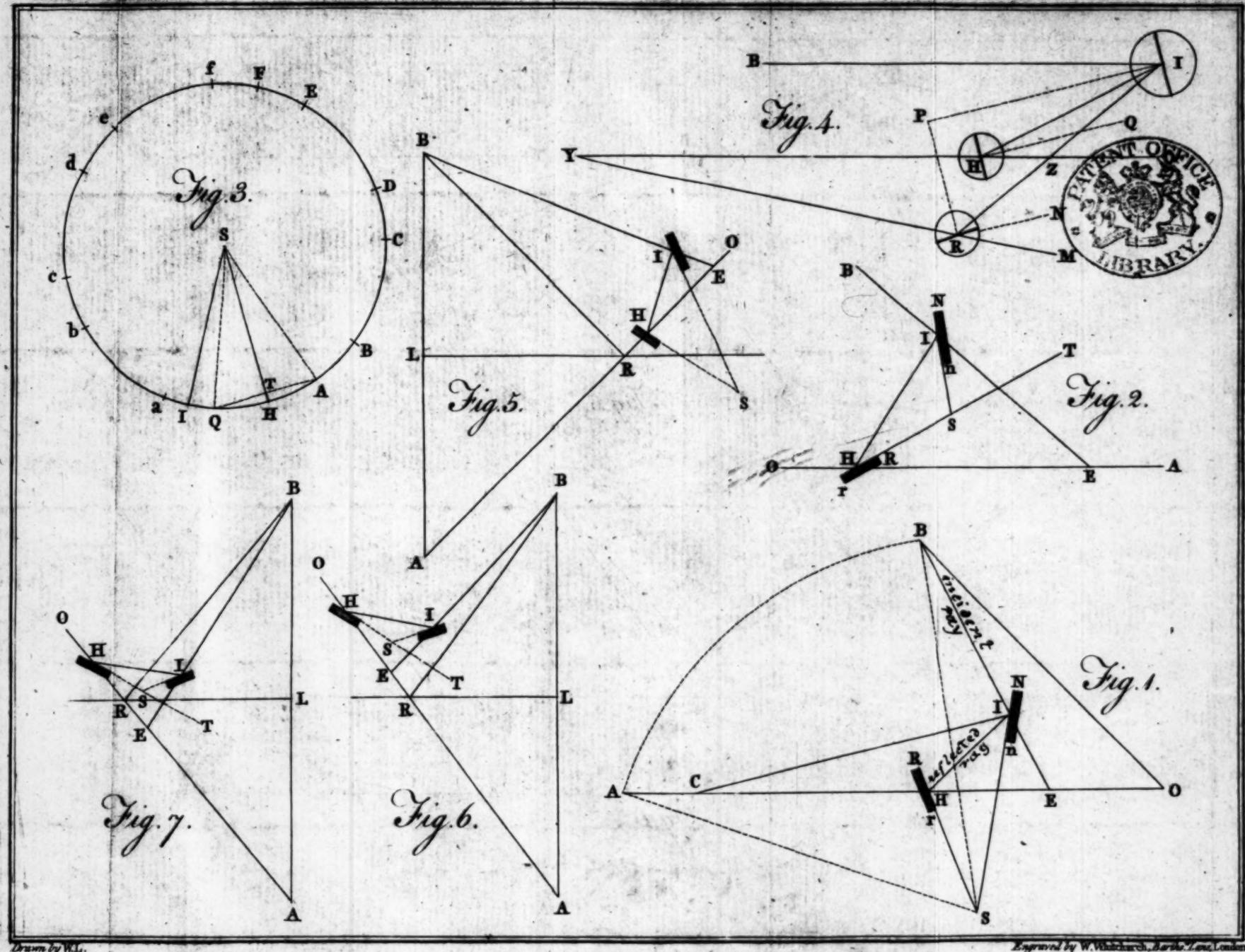
* See the *American Transactions*, Vol. I. printed at Philadelphia, 1771, Appendix, page 21.

T H E E N D.









Drawn by W.L.

Engraved by W. H. Worrell.

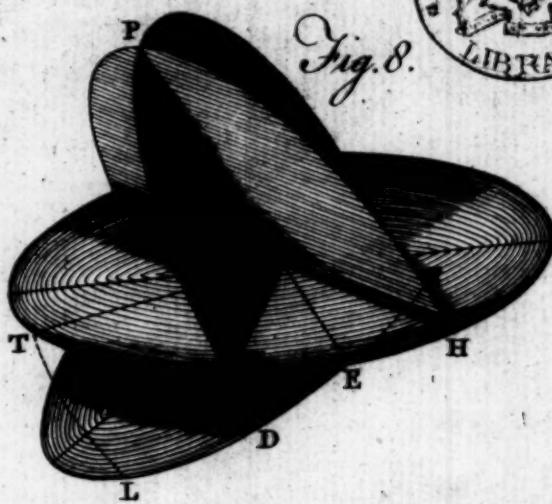


Fig. 8.

Fig. 10.

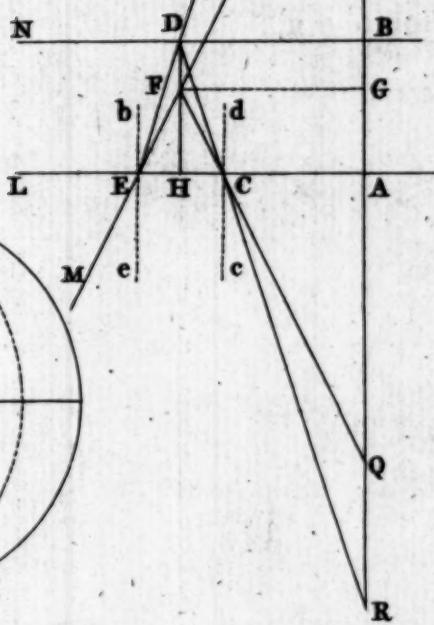


Fig. 9.

